

The Role of Western Boundaries in Wind-Driven Energetics

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EOAS

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with Q Jamet, B Deremble and N Wienders

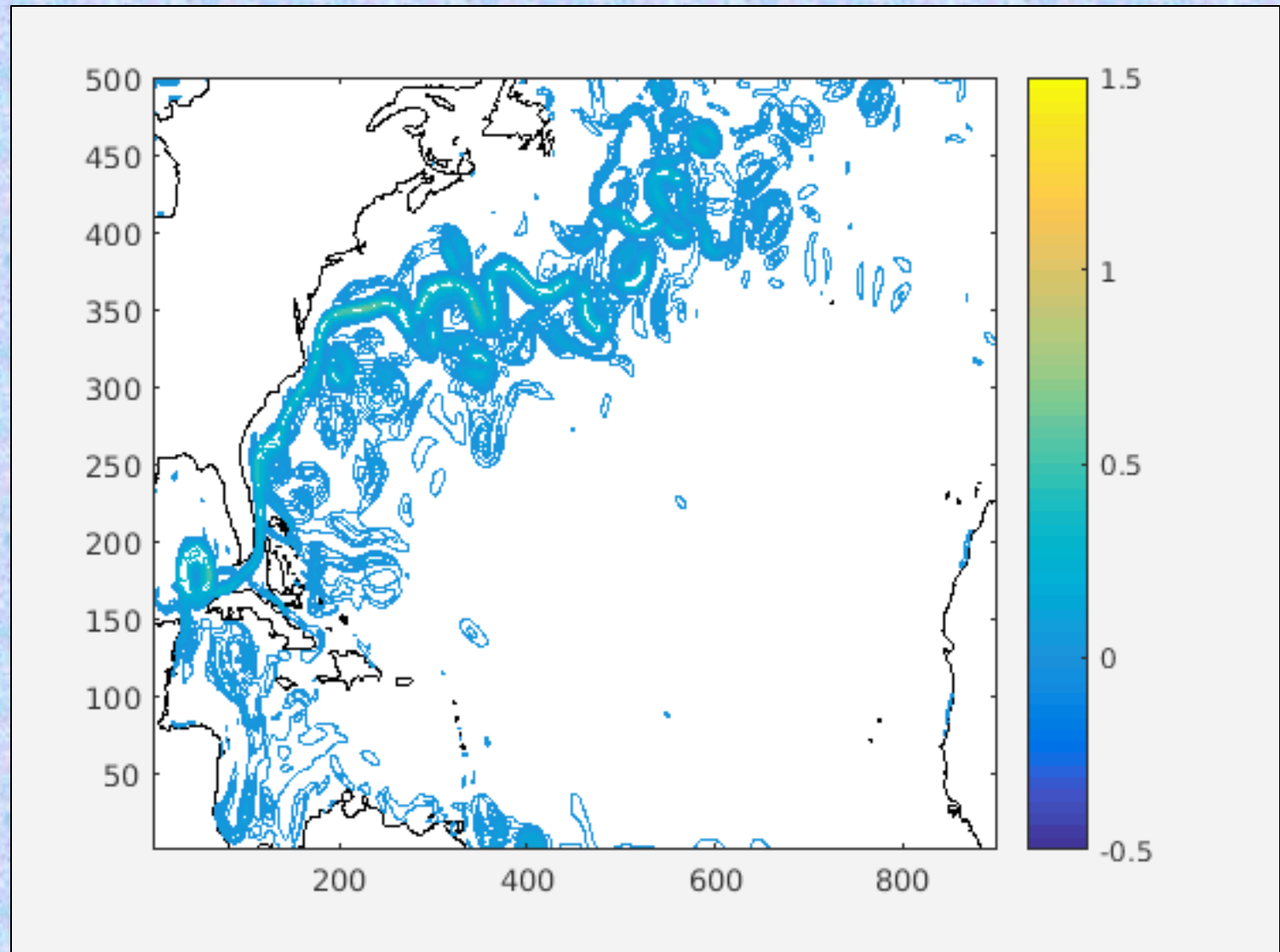
Sources and Sinks of Ocean Mesoscale Eddy Energy

Tallahassee

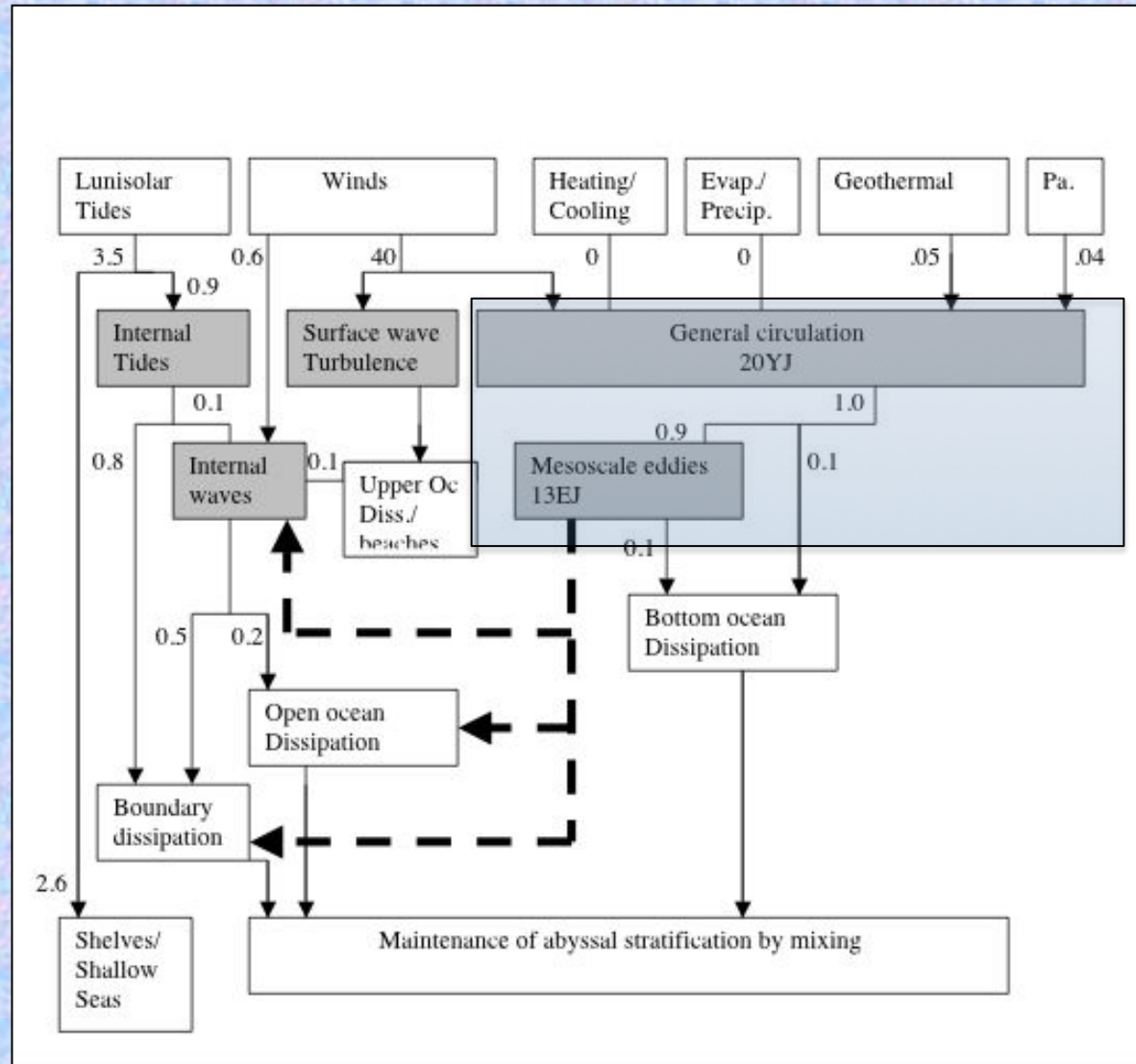
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Surface Speed (m/s) from 1/12 North Atlantic Model



A recent view of the Ocean Energy Budget



The leading theoretical views of the General Circulation and eddies come from

$$\beta v = f \frac{\partial}{\partial z} w; \quad \vec{u} \cdot \nabla f N^2 = 0$$

$$\varepsilon = \frac{U}{fL} \ll 1, \quad L \sim L_\beta, \quad \frac{\delta h}{H} \sim 1$$

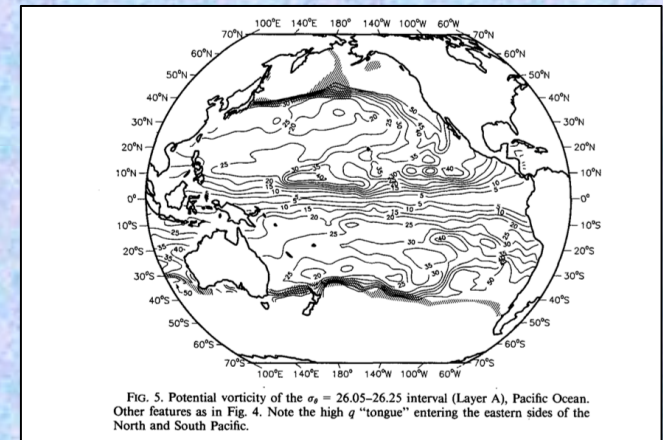
Ventilated Thermocline Theory
geostrophic, hydrostatic, steady

$$\frac{d}{dt} q = 0; \quad q = \nabla^2 p + \frac{\partial}{\partial z} \frac{f}{N^2} \frac{\partial}{\partial z} p + \beta y$$

$$\varepsilon = \frac{U}{fL} \ll 1, \quad L \sim L_R, \quad \frac{\delta h}{H} \sim \varepsilon$$

Homogenization Theory
mildly ageostrophic

Many Studies have supported the explanatory value of these theories to the ocean.



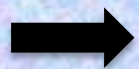
Space and time scale separation

$$T_{pg} = \frac{L_\beta}{\beta L_R^2} \gg \frac{1}{\varepsilon f}$$

$$L_\beta \gg L_R$$

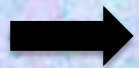
Pedlosky 1984

The Equations for Geostrophic Motion in the Ocean



$$\frac{\partial}{\partial T} \frac{(z_o)_b}{f} + \frac{1}{f} \bar{J}(M_o, \frac{(z_o)_b}{f}) = 0$$

The Ventilated Thermocline



$$\frac{\partial}{\partial t} q + \underline{\bar{u}_o} \cdot \nabla q + J(M_1, q) + \frac{(z_o)_b}{f} J(M_1, \underline{\frac{f}{(z_o)_b}}) = 0$$

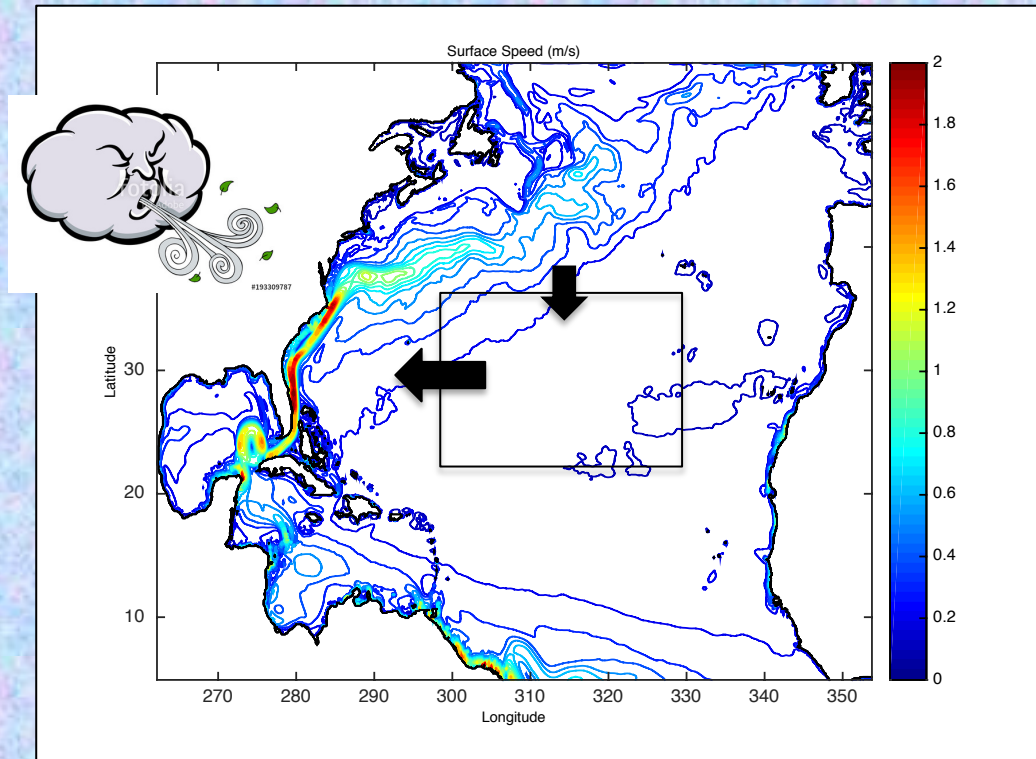
Quasi-geostrophy

$$q = \nabla^2 M_1 - f \frac{(M_1)_{bb}}{(z_o)_b}$$

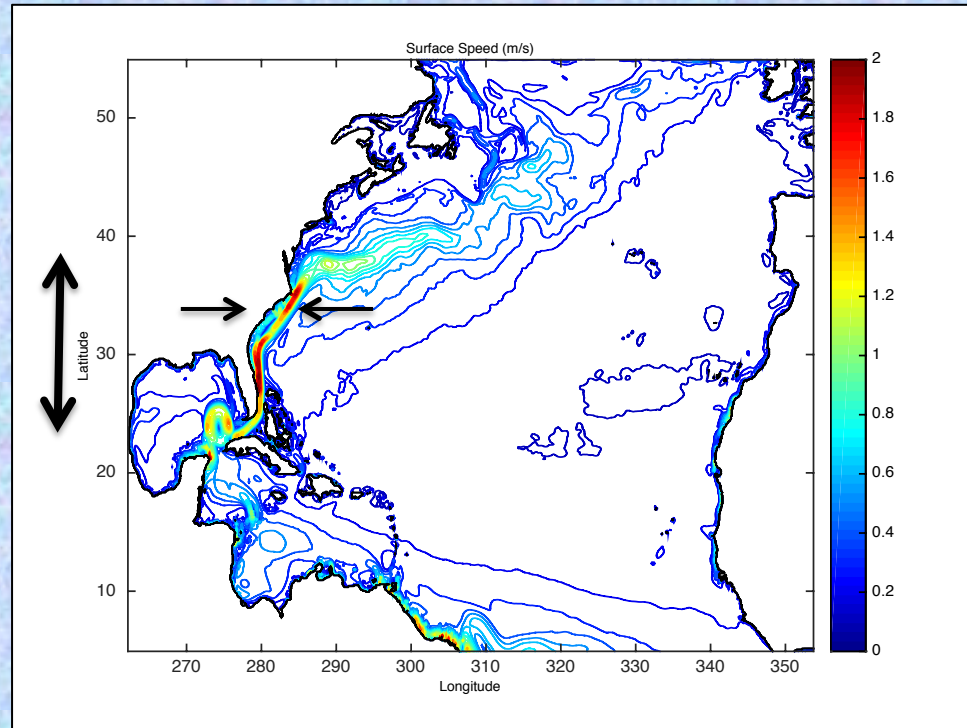
Looks really good - BUT

Leaves us with an energetics problem
Ventilated Thermocline is forced, but not
dissipative

$$\iint \vec{u}(p - bz) \cdot \vec{n} dS = \iint_A \vec{\tau} \cdot \vec{u}_o dA$$



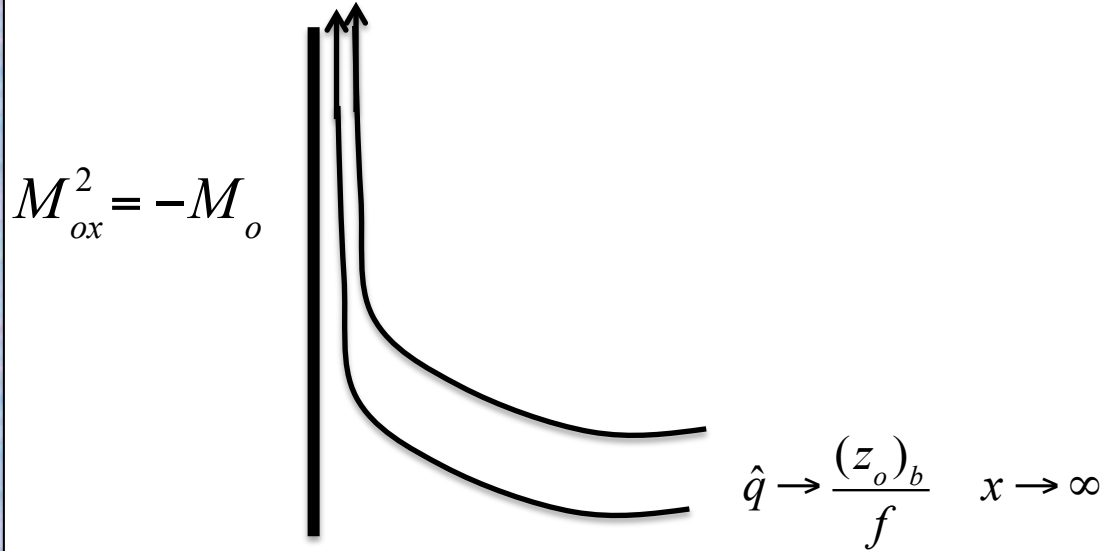
It's natural to look to western boundary layers for a solution



Instead of $X, Y \sim L_\beta$, $x, y \sim L_R$
Examine $x \sim L_R$, $y \sim L_\beta$

Grooms, et al DAO, 2011

After much sweating and grinding of teeth, bl
 pv equation emerges that connects to the pg
 interior



$$u_o \hat{q}_x + \hat{v}_o \hat{q}_y = 0;$$

$$\hat{q} = \frac{\frac{M_{ox}}{f} + f}{(M_o)_{bb}}$$

Looks really good - BUT

Leaves us with an energetics problem
Energy is conservative.



$$\bar{u}(-bz + p)$$

$$\bar{u}K$$

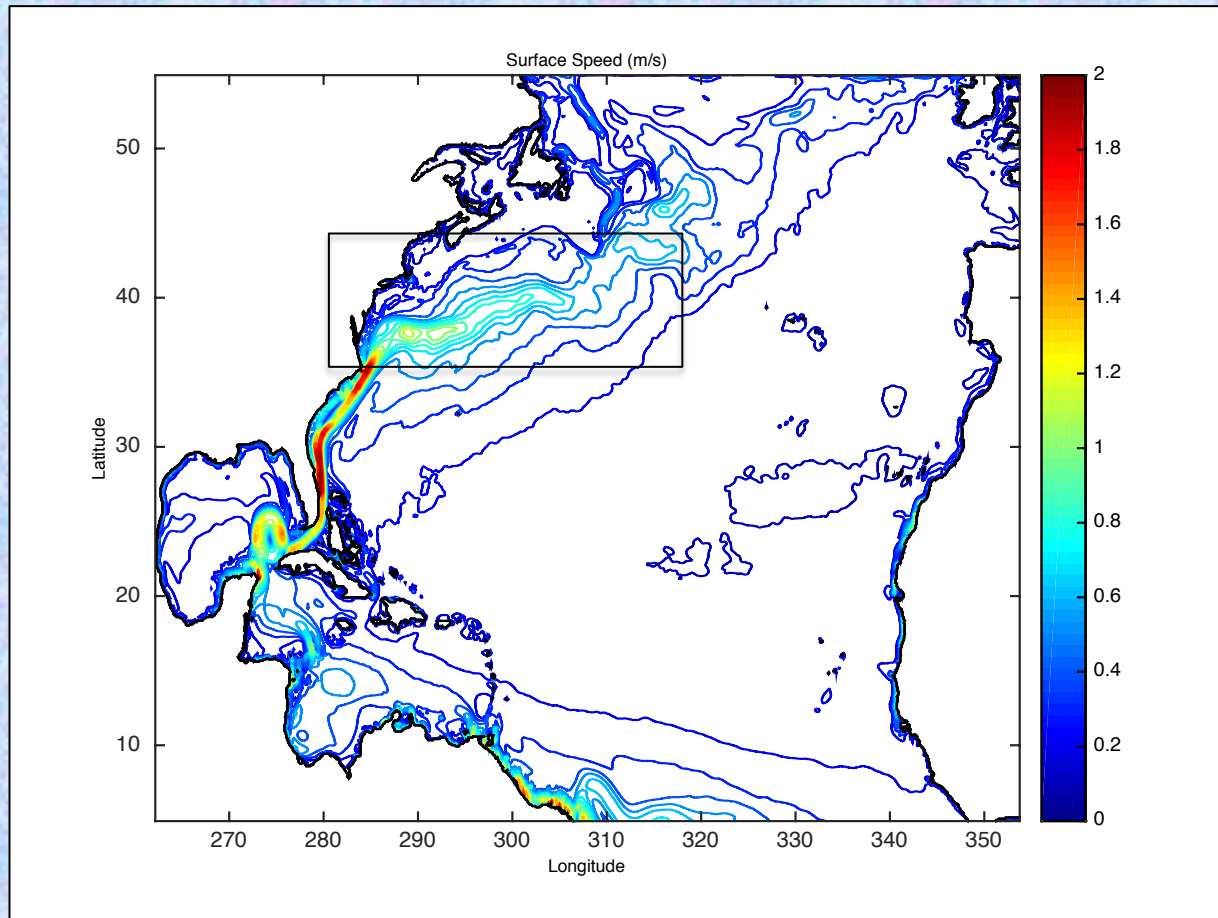


$$\iint_S \bar{u}(K - bz + p) \cdot \vec{n} dS = 0$$

Both boundary
conditions are steady

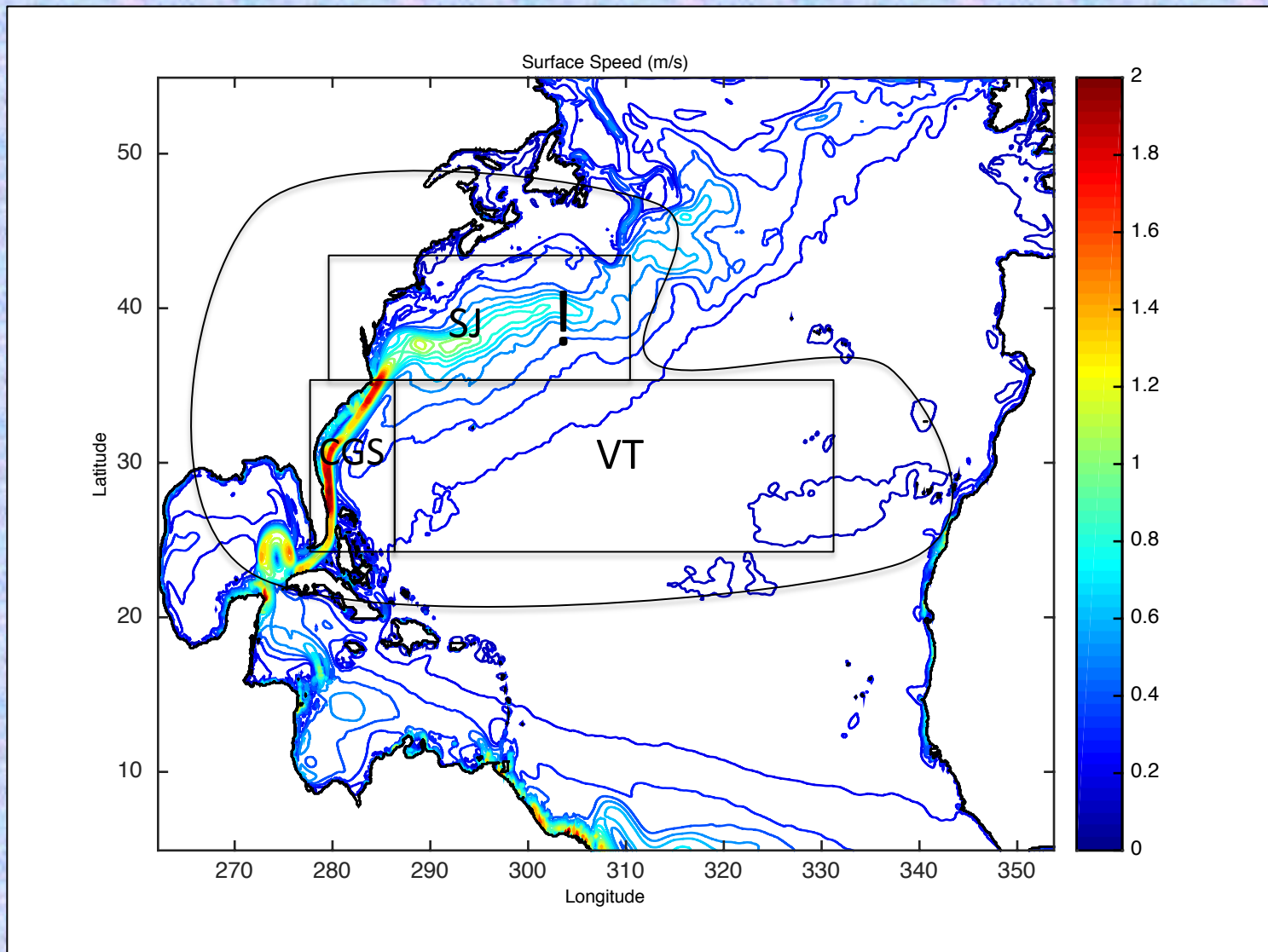
So – What's Next?

Continuing around the gyre, we come to the separated jet



Full PE
dynamics
are
required

So, where does this leave us?



How to test in a realistic GCM?

$$\vec{u} \cdot \nabla \vec{u}_h + \vec{f}x\vec{u}_h = -\nabla_h \bar{p} - \nabla \cdot \vec{F} - \nabla \cdot \overline{\vec{u}'\vec{u}'_h}$$



$$\nabla \cdot \vec{u}K = -\overline{\vec{u}_h \cdot \nabla \bar{p}} - \vec{u}_h \cdot \nabla \cdot \vec{F} - \vec{u}_h \nabla \cdot \overline{\vec{u}'\vec{u}'_h}$$

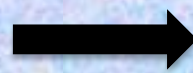
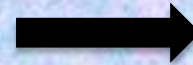
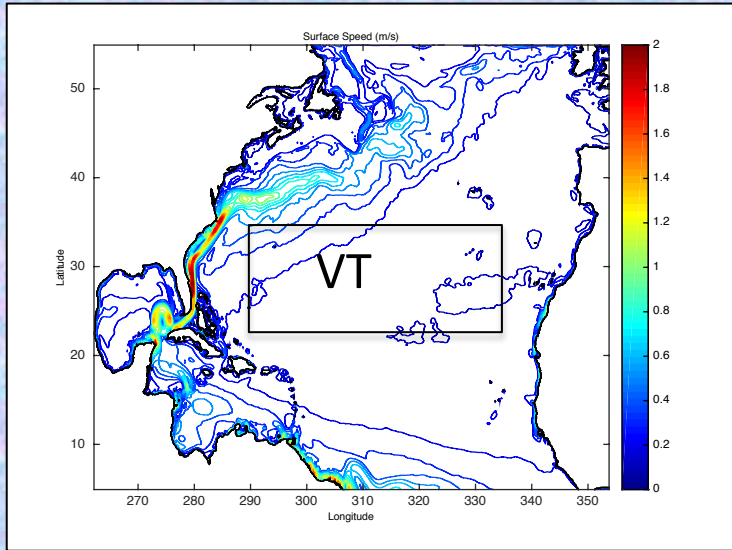
↑
MKEF

↑
PW

↑
VIS

↑
MEC

WW+KEDISS



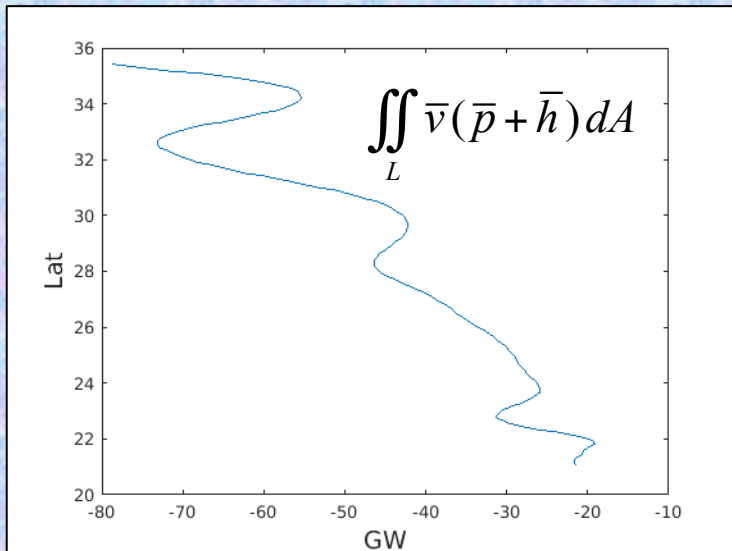
DKEF = .01 GW

PW = 3.5 GW

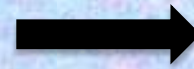
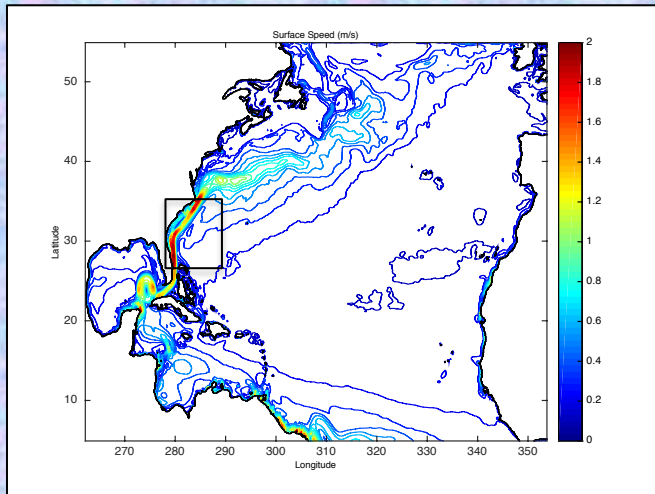
WW = 4.9 GW

KEDISS = .01 GW

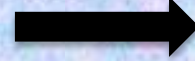
MEC = -1.3 GW



Something of a surprise



$$\text{DKEF} = 21 \text{ GW}$$



$$\text{PW} = -39 \text{ GW}$$

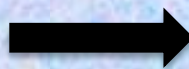
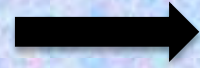
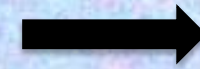
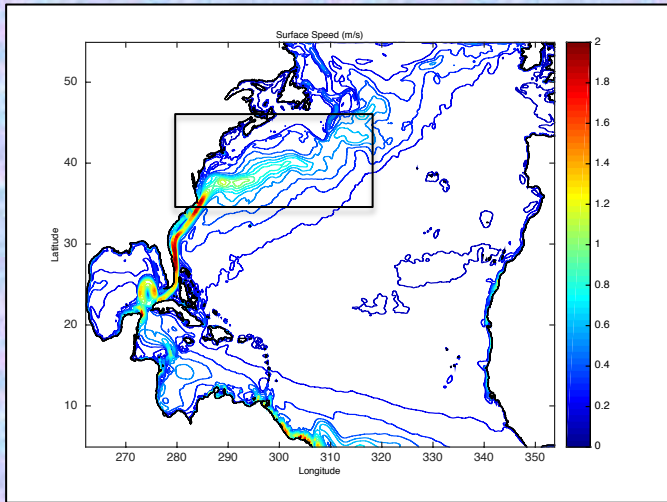
$$\text{WW} = -0.7 \text{ GW}$$

$$\text{KEDISS} = -7 \text{ GW}$$

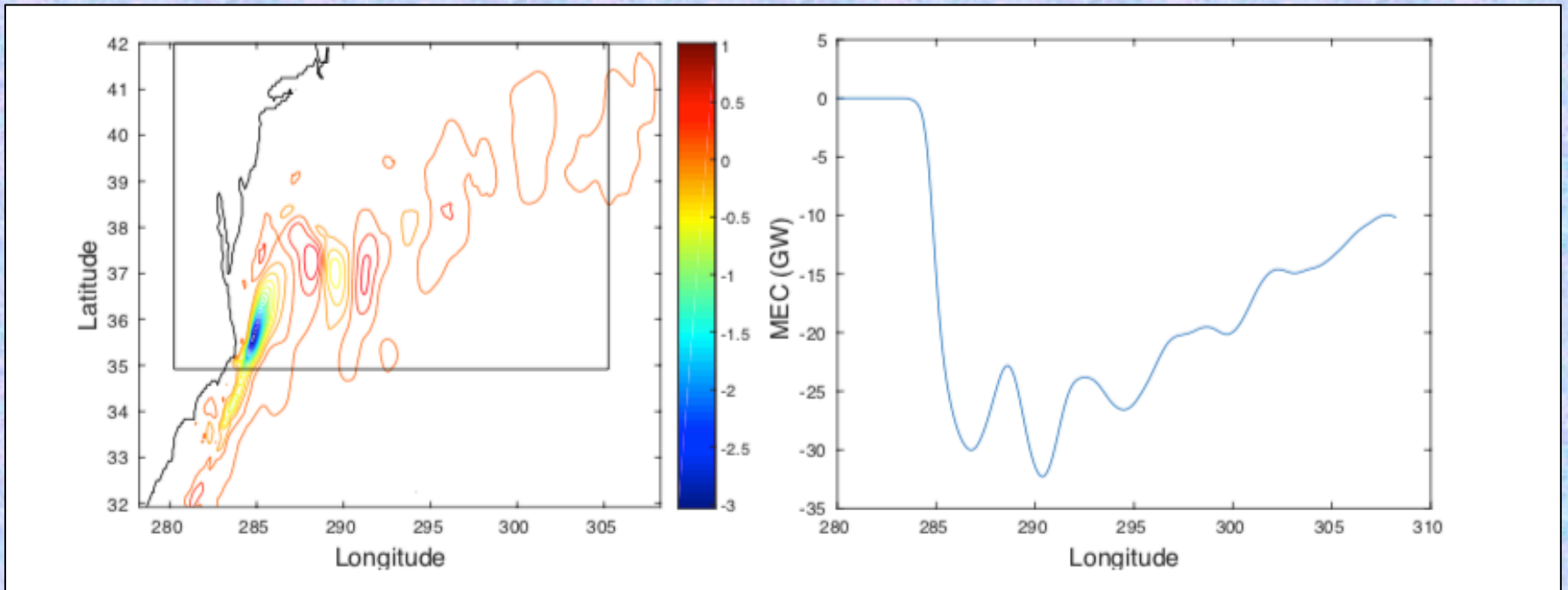
$$\text{MEC} = -11 \text{ GW}$$

Leading order balance -
Potential energy to kinetic
Energy conversion

Dissipation and eddies are not negligible, the latter due to Charleston Bump



DKEF = -28 GW
PW = 17 GW
WW = 3.6 GW
KEDISS = -0.7 GW
MEC = -14 GW



Summary:

A straightforward dynamical division of subtropical gyres suggested by theory seems consistent with a realistic 1/12 model

Mesoscale is powered up primarily in the separated jet extension (not obviously qg)

Is the ventilated thermocline inertial?