

From Virtual Navier-Stokes Flows to Numerical Atmosphere & Ocean Models or *A Discrete Model Hierarchy*

Peter Korn

Max Planck Institute for Meteorology, Hamburg

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I will discuss the following topics

- i Incompressible Dynamics (*~ ocean*)
- ii Compressible Dynamics (*~ atmosphere*)
- iii Singular Limits (*relation between different equations*)
- iv Lessons learned

Focus on **finite-dimensional** setup and **nonhydrostatic dynamics**

Starting Point: Primitive Equations - Hydrostatic and Boussinesq

Velocity field: $\mathbf{v} = (v_h, w)$, horizontal velocity v_h , vertical velocity w

$$\partial_t \mathbf{v}_h + \omega_z \vec{e}_z \times \mathbf{v}_h + \frac{\nabla_h |\mathbf{v}_h|^2}{2} + w \partial_z \mathbf{v}_h + \frac{1}{\rho_0} \nabla_h \mathbf{p} - \mathcal{D} \mathbf{v}_h = 0$$

$$\partial_z \mathbf{p} = -\rho \mathbf{g}$$

$$\partial_t \eta + \operatorname{div}_h \int_{-B}^{\eta} \mathbf{v} \, dz = 0$$

$$\operatorname{div}_h \mathbf{v}_h + \partial_z w = 0$$

$$\partial_t C + \operatorname{div}(C \mathbf{v}) - \operatorname{div}(\mathbb{K}^C \nabla C) = 0$$

$$\rho = F_{eos}(\rho, T, S),$$

Starting Point: Primitive Equations - Hydrostatic and Boussinesq

$$\partial_t \mathbf{v}_h + \omega_z \vec{\mathbf{e}}_z \times \mathbf{v}_h + \frac{\nabla_h |h\mathbf{v}|^2}{2} + w \partial_z \mathbf{v}_h + \frac{1}{\rho_0} \nabla_h p - \mathcal{D} \mathbf{v}_h = 0$$

$$\partial_z p = 0 \quad \text{--- } \rho g$$

$$\partial_t \eta + \operatorname{div}_h \int_{-B}^{\eta} v \, dz = 0$$

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$$\partial_z p = 0 - \cancel{\rho g}$$

$$\cancel{\partial_t \eta + \operatorname{div}_h \int_{-B}^{\eta} v \, dz = 0}$$

$$\operatorname{div}_h \mathbf{v}_h + \partial_z w = 0$$

$$\cancel{\partial_t C + \operatorname{div}(C\mathbf{v}) - \operatorname{div}(\mathbb{K}^C \nabla C) = 0}$$

$$\cancel{\rho = F_{\text{eos}}(p, T, S),}$$

*This is the Hydrostatic Euler Equation.
How can we make it NonHydrostatic ?*

Route A to Nonhydrostatic Euler: add w -eq to hydrostatic eqs.

$$\partial_t \mathbf{v}_h + \omega_z \vec{\mathbf{e}}_z \times \mathbf{v}_h + w \partial_z \mathbf{v}_h + \nabla_h \left(p + \frac{|\mathbf{v}_h|^2}{2} \right) = 0,$$

$$\partial_t \mathbf{w} + (\mathbf{v}, \mathbf{w}) \cdot \nabla \mathbf{w} + \partial_z p = 0,$$

Route B to Nonhydrostatic Euler: 3D vector-invariant

$$\partial_t \mathbf{v} + \boldsymbol{\omega} \times \mathbf{v} + \nabla \left(p + \frac{|\mathbf{v}|^2}{2} \right) = 0 \quad (\mathbf{v}, \boldsymbol{\omega} \text{ are 3D vector fields})$$

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$$\partial_t w + (\mathbf{v}, w) \cdot \nabla w + \partial_z p = 0,$$

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- **Route A:** easy to implement, breaks beauty of Euler equation. (*no consistent vorticity eq., energetics presumably impossible...*)
- **Route B** challenge is discrete exterior product $\boldsymbol{\omega} \times \mathbf{v}$

Strategy: we go for Route B

- ① Focus on inviscid case and get conservation properties
- ② Incorporate dissipation via *explicit dissipation, upwind-biased ...*

Incompressible Euler Equations

$$\partial_t \mathbf{v} + \boldsymbol{\omega} \times \mathbf{v} + \nabla \left(p + \frac{|\mathbf{v}|^2}{2} \right) = 0, \quad \operatorname{div} \mathbf{v} = 0$$

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Advection - Continuous Cross Product

$$\text{horiz. v-equation: } (\boldsymbol{\omega} \times \mathbf{v})|_h = \begin{pmatrix} \omega_y v_z - \omega_z v_y \\ \omega_z v_x - \omega_x v_z \end{pmatrix},$$

$$\text{vert. v-equation: } (\boldsymbol{\omega} \times \mathbf{v})|_v = \omega_h \cdot \mathbf{v}_h^\perp = (\omega_x v_y - \omega_y v_x).$$

- **blue/Hydrostatic:** Cross-product terms with vertical vorticity ω_z
- **red/Nonhydrostatic:** Cross-product terms with horiz. vorticity ω_h
- \rightarrow we have **blue** we need **red**

We need 3D vorticity vector (ω_h, ω_z) and construct missing ω_h via Stokes Theorem

Horizontal component of vorticity vector - continuous

$$\omega_h := \text{curl}_h \mathbf{v} = \begin{pmatrix} \omega_x \\ \omega_y \end{pmatrix} = \begin{pmatrix} \partial_y v_z - \partial_z v_y \\ \partial_z v_x - \partial_x v_z \end{pmatrix}$$

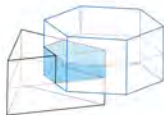
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Three Observations

- 1 Prismatic grid: 2D horizontal \times 1D vertical
- 2 Dual prism is shifted horizontally and vertically
- 3 Vertical faces are rectangles !

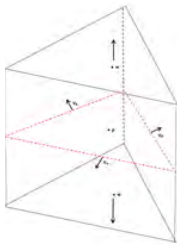
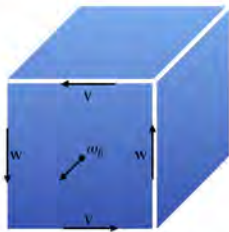


Horizontal component of vorticity vector $\omega_{\partial P}$ via Stokes

$$\omega_h \underbrace{\quad \rightarrow}_{\text{by Stokes}} \quad \mathbf{curl}_h U_{\partial P} := w_{K,k+1/2} + v_{e,k} - w_{L,k+1/2} - v_{e,k+1}$$

This defines

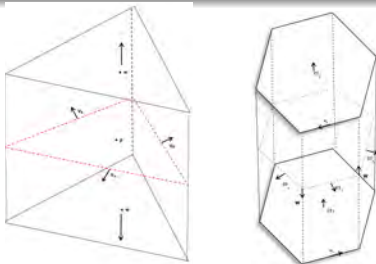
- discrete 3D curl-operator: $\mathbf{curl} U_{\partial P} = (\mathbf{curl}_h U_{\partial P}, \mathbf{curl}_v U_{\partial P})$
- vorticity vector at faces of dual prism $\omega_{\partial P} := (\omega_h, \omega_v)$



$$\begin{pmatrix} \omega_y v_z - \omega_z v_y \\ \omega_z v_x - \omega_x v_z \\ \omega_y v_y - \omega_y v_x \end{pmatrix} \rightsquigarrow \omega_{\partial \hat{\mathcal{P}}} \star u_{\partial \mathcal{P}} := \begin{pmatrix} \hat{\mathcal{P}}_h^\dagger(\omega_z \hat{\mathcal{P}}_h v) - \mathcal{P}_z \mathcal{P}^T (w \tilde{\mathcal{P}}_h \omega_h) \\ \tilde{\mathcal{P}}_h \omega_h \cdot \mathcal{P} \mathcal{P}_z^T v \end{pmatrix}$$

the secret sauce

- $\mathcal{P}, \hat{\mathcal{P}}, \tilde{\mathcal{P}}$ are Hilbert space compatible reconstructions



Kernel of Differential Operators

- i **grad** $p = 0$ if and only if p is constant
- ii **curl** $v = 0$ if and only if $v = \mathbf{grad} p$
- iii **div** $v = 0$ if and only if $v = \mathbf{curl}^T u$.

Discrete Biot-Savart:

From given $\omega \in \mathcal{H}_{\hat{V}}$ the velocity $u_{\partial P} \in \mathcal{H}_{\partial P}$ with $\mathbf{div} \mathcal{M} u_{\partial P} = 0$, is recovered by solving Laplace equation

$$\mathbf{div} \mathcal{M} \mathbf{grad} u_{\partial P} = \mathbf{curl}^T \omega.$$

Cont. Velocity Space & Discrete Degrees of Freedom on Prism Q

- $\mathbb{F}(Q) := \{f \in H_{div}(Q) \cap H_{rot}(Q) : \operatorname{div} f \in \mathbb{P}_0(Q), \operatorname{curl} f = 0, f|_e \cdot \mathbf{n}_e \in \mathbb{P}_0(Q) \forall e \in \partial Q\},$
- $\operatorname{dof}_{\mathbb{F}(Q)}(f) := \Pi f := \frac{1}{|e|} \int_e f \cdot \mathbf{n}_e \, ds, \quad \forall e \in \partial Q.$

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Theorem

Discrete DoF's above are unisolvent, i.e. they characterize uniquely the respective continuous virtual element space.

Reconstructions of disc. DoF via local PDEs

Given discrete velocity dof's $v_e \in \partial Q$. Define continuous function $\tilde{v} := \mathcal{P}v$ on Q as solution of local div-curl problem

$$\begin{aligned} \operatorname{div} \tilde{v} &= \mathbf{div} v, & \text{on } Q, \\ \operatorname{curl} \tilde{v} &= 0, & \text{on } Q, \\ \tilde{v} \cdot \mathbf{n}_e &= v_e & \text{on } \partial Q. \end{aligned}$$

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Scalar Product on Discrete Velocity Space in Terms of Reconstructions

$$\begin{aligned} \langle u, v \rangle_{\mathbb{F}(Q)} &:= \int_Q \mathcal{P}u \cdot \mathcal{P}v \, dx \\ \int_{\Omega} \mathcal{P}u \cdot \mathcal{P}v \, dx &= \sum_{Q \in \mathcal{C}} |Q| \mathcal{P}u_Q \cdot \mathcal{P}v_Q, \end{aligned}$$

Pressure & Vorticity

For pressure and vorticity spaces similar scalar products via reconstructions via local div-curl PDEs.

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Fundamental Lemma on Reconstructions

Let $\mathcal{P} : v_e \rightarrow \mathcal{P}v \in \mathbb{F}(Q)$ be a reconstruction such that

- \mathcal{P} is the right-inverse of projection $\Pi f := \frac{1}{|e|} \int_e f \cdot \mathbf{n}_e d$
- \mathcal{P} is first-order accurate
- \mathcal{P} commutes with continuous differential operators *grad*, *div*, *curl*
- Reconstructed functions are orthogonal to linear polynomials on Q with zero mean
- \mathcal{P} has a local stencil

Then it holds $\int_Q \mathcal{P}v \cdot \mathbf{e}_i dx = \sum_{e \in \partial Q} v_e |e| (\mathbf{x}_e - \mathbf{x}_Q) \cdot \mathbf{e}_i$.

(Analogous results for $\hat{\mathcal{P}}, \tilde{\mathcal{P}}$)

Reconstructions

- Div-Curl- PDE are actually never solved.
- Reconstructions have an explicit & computable form.
- We need three Reconstructions
 - \mathcal{P} : face dof \rightarrow inside primal 3D prism
 - $\hat{\mathcal{P}}$: face dof \rightarrow inside 3D dual prism
 - $\tilde{\mathcal{P}}_h$: edge dof \rightarrow inside 2D primal cell
 - $\mathcal{M} := \mathcal{P}^T \mathcal{P}$

End of Numerical Disgression - Back to Euler Equations

Incompressible Euler

- $$\bullet \left\langle \frac{d}{dt} \mathcal{M} u_{\partial P}, \phi \right\rangle_{\mathcal{H}_{\partial P}} + \left\langle \omega_{\partial P} \star u_{\partial P}, \phi \right\rangle_{\mathcal{H}_{\partial P}}$$

$$+ \left\langle \mathcal{M} \mathbf{grad} \left(p + \frac{|\mathcal{P} u_{\partial P}|_{\mathbb{R}^3}^2}{2} \right), \phi \right\rangle_{\mathcal{H}_{\partial P}} = 0, \quad \forall \phi \in \mathcal{H}_{\partial P},$$
- $$\bullet \mathbf{div} \mathcal{M} u_{\partial P} = 0.$$

Incompressible Euler

- $\left\langle \frac{d}{dt} \mathcal{M}u_{\partial P}, \phi \right\rangle_{\mathcal{H}_{\partial P}} + \left\langle \omega_{\partial \hat{P}} \star u_{\partial P}, \phi \right\rangle_{\mathcal{H}_{\partial P}}$
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- $\mathbf{div} \mathcal{M}u_{\partial P} = 0.$

Pressure Recovery

Pressure is recovered from $u_{\partial P} \in \mathcal{H}_{\partial P}$ by solving Laplace equation

$$\mathbf{div} \mathcal{M} \mathbf{grad} p = -\mathbf{div} \mathcal{M} \mathbf{grad} \left(\frac{|\mathcal{P}_3 u_{\partial P}|_{\mathbb{R}^3}^2}{2} \right) - \mathbf{div} \mathcal{L}(\omega_{\partial \hat{P}}, u_{\partial P})$$

Theorem (Well-Posedness of Semi-Discrete Euler Equations)

- *Let a time interval $[0, T]$ and initial conditions $u_0 \in \mathcal{H}_{\partial P}^{\text{div}}$ be given.*

Then there exist for $t \in [0, T]$ a unique solution $u_{\partial P}(t) \in \mathcal{H}_{\partial P}$ of the discrete Euler equations.

(Proof by Picard's theorem for ODE for short time and extension to long time via energy conservation.)

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Well-Posedness of Semi-Discrete Navier-Stokes

Proof translates to Navier-Stokes equations,
with dissipation given by $\mathcal{D}(v) := \mathbf{curl}^T(\nu \mathbf{curl} u_{\partial P})$.

Linear Momentum

Let $u_{\partial P} \in \mathcal{H}_{\partial P}^{\text{div}M}$ be a solution of the Euler equation. Then the linear momentum $\mathcal{I} := \langle \mathcal{P}^T \mathcal{P} u_{\partial P}, 1 \rangle_{\mathbf{H}_{\mathcal{P}}}$ satisfies

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Angular Momentum $L := \vec{x} \times \vec{u}$

- Question: How to define $L := \vec{x} \times \vec{u}$ for staggered $\vec{u} = (v_h, w)$?
- Observation: Vector product $\vec{\omega} \times \vec{u} \sim \omega_{\partial \hat{P}} \star u_{\partial P}$

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Angular Momentum - Definition and Conservation

Define discrete angular momentum:

$$\ell(u_{\partial P}) := \vec{x} \star u_{\partial P},$$

where \vec{x} is coordinate vector of the position of vorticity at dual prism faces. Then

$$\frac{d}{dt} \langle \ell(u_{\partial P}), 1 \rangle = 0$$

Theorem (Energy Conservation)

The solution $u_{\partial P}(t) \in \mathcal{H}_{\partial P}^{\text{div}}$ of the discrete incompressible Euler equations conserves kinetic energy:

$$\frac{d}{dt} E^{\text{kin}}(t) = 0, \quad E^{\text{kin}}(t) := \|\mathcal{P}_3 u_{\partial P}(t)\|_{\mathbf{H}^{\mathcal{P}}}^2$$

2D: *Energy and enstrophy conservation.*

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2D: Energy and enstrophy conservation.

Helicity: Inner Product of Velocity & Vorticity

$$\mathcal{H} := \int_{\Omega} \mathbf{v} \cdot \omega \, dx \quad | \quad \mathbf{H} := \langle \hat{\mathcal{P}} \mathbf{v} \omega_z, \mathbf{1} \rangle_{\mathcal{H}_{\hat{\mathbf{v}}}} + \langle \mathbf{w} \tilde{\mathcal{P}}_h \omega_h, \mathbf{1} \rangle_{\mathcal{H}_P}$$

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Theorem (Helicity Conservation)

The solution $u_{\partial P_k}(t) \in \mathcal{H}_{\partial P}^{\text{div}}$ of the discrete **3D** incompressible Euler equations conserves helicity:

$$d_t \mathbf{H} = 0.$$

(Proof by combining equations for vorticity and velocity and suitable test functions in discrete weak form.)

Time stepping - Implicit

$$\begin{aligned}
& \left\langle \frac{\mathcal{M}(u_{\partial P}^{n+1} - u_{\partial P}^n)}{\Delta t}, \phi \right\rangle_{\mathcal{H}_{\partial P}} + \left\langle \omega_{\partial \hat{P}}^{n+1/2} \star u_{\partial P}^{n+1/2}, \phi \right\rangle_{\mathcal{H}_{\partial P}} \\
& + \left\langle \mathcal{M} \mathbf{grad} \left(p^{n+1/2} + \frac{|\mathcal{P} u_{\partial P}^{n+1/2}|_{\mathbb{R}^3}^2}{2} \right), \phi \right\rangle_{\mathcal{H}_{\partial P}} = \left\langle f^{n+1/2}, \phi \right\rangle_{\mathcal{H}_{\partial P}}, \\
& \mathbf{div} \mathcal{M} u_{\partial P}^{n+1} = 0,
\end{aligned}$$

where $u_{\partial P}^{n+1/2} := \frac{1}{2}(u_{\partial P}^{n+1} + u_{\partial P}^n)$

Theorem

• Let $\Delta t > 0$ be the time step size and $u_{\partial P}^0$ in $\mathcal{H}_{\partial P}^{\text{div}M}$ initial conditions. Then a unique solution exists $(u_{\partial P}^n, p^n)$ of incompressible Euler with the following properties:

- ① $(u_{\partial P}^n, p^n)$ conserves global kinetic energy, $E^{\text{kin}}(u_{\partial P}^n) = E^{\text{kin}}(u_{\partial P}^0)$
- ② $(u_{\partial P}^n, p^n)$ conserves linear momentum $\mathcal{I}(u_{\partial P}^n) = \mathcal{I}(u_{\partial P}^0)$ and angular momentum $\ell(u_{\partial P}^n) = \ell(u_{\partial P}^0)$
- ③ $(u_{\partial P}^n, p^n)$ conserves vorticity $\langle \omega_{\partial \hat{P}}^n, 1 \rangle_{\mathcal{H}_{\hat{V}}} = \langle \omega_{\partial \hat{P}}^0, 1 \rangle_{\mathcal{H}_{\hat{V}}}$.
- ④ $(u_{\partial P}^n, p^n)$ conserves helicity $\mathbf{H}(u_{\partial P}^n) = \mathbf{H}(u_{\partial P}^0)$
- ⑤ $(u_{\partial P}^n, p^n)$ is reversible in time

(Proof: Schauder fix point theorem, differentiability of mapping for uniqueness. Conservation properties rely on implicit time stepping.)

Navier-Stokes Equations

Proof applies to Navier-Stokes, without conservation props.

Now allow the density to vary - but not to compress

Incompressible Euler-Boussinesq Equations

$$\partial_t \mathbf{v} + \omega \times \mathbf{v} + \nabla \left(p + \frac{|\mathbf{v}|^2}{2} \right) = g\rho \vec{e}_z, \quad \text{div} \mathbf{v} = 0$$

$$\partial_t \rho + \text{div}(\rho \mathbf{v}) = 0.$$

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Euler-Boussinesq

- $\left\langle \frac{d}{dt} \mathcal{M} u_{\partial P}, \phi \right\rangle_{\mathcal{H}_{\partial P}} + \left\langle \omega_{\partial P} \star u_{\partial P}, \phi \right\rangle_{\mathcal{H}_{\partial P}}$
 $+ \left\langle \mathcal{M} \operatorname{grad} \left(p + \frac{|\mathcal{P} u_{\partial P}|_{\mathbb{R}^3}^2}{2} \right), \phi \right\rangle_{\mathcal{H}_{\partial P}} = \left\langle g\rho \vec{e}_z, \phi \right\rangle_{\mathcal{H}_P}$
- $\operatorname{div} \mathcal{M} u_{\partial P} = 0,$
- $\left\langle \partial_t \rho, \psi \right\rangle_{\mathcal{H}_P} + \left\langle \operatorname{div}(\mathcal{P}^T(\rho \mathcal{P} \mathbf{v})), \psi \right\rangle_{\mathcal{H}_P} = 0.$

Theorem (Well-Posedness of Semi-Discrete Euler-Boussinesq)

i) A unique solution $u_{\partial P}(t) \in \mathcal{H}_{\partial P}^{div}$, $\rho \in \mathcal{H}_P$ to discrete Euler-Boussinesq equations exists

ii) and it has the following properties

- **Energy Conservation:** The sum of kinetic and potential energy is conserved

$$\frac{d}{dt}(E^{kin} + E^{pot})(t) = 0, \quad E^{pot} := g\rho Qw$$

- **Helicity Conservation** The solution $u_{\partial P_k}(t) \in \mathcal{H}_{\partial P}^{div}$ of the discrete **3D** incompressible Euler equations satisfies

$$d_t \mathbf{H} = F(\Phi). \quad (\Phi \text{ geopotential})$$

PV - Continuous:

$$\mathcal{PV} := \omega \cdot \nabla \rho$$

PV - Discrete: Inner Product of ω and $\mathbf{grad} \rho$

$$\mathbf{PV}(u_{\partial P})|_{\hat{P}} := \langle \omega_z \hat{P}_h \mathbf{grad} \rho, \mathbf{1} \rangle_{\mathbf{H}_{\partial \hat{P}}} + \langle \omega_h \mathbf{D}_z \rho, \mathbf{1} \rangle_{\mathcal{H}_{\partial \hat{P}}}$$

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Potential Vorticity Conservation

Let $u_{\partial P} \in \mathcal{H}_{\partial P}^{divM}$ be a solution of the Euler-Boussinesq equations. Then

$$\frac{d}{dt} \langle \mathbf{PV}, \mathbf{1} \rangle_{\mathcal{H}_{\hat{P}}} = 0.$$

Proof by combining equations for vorticity and for density gradient.

Theorem - Implicit Time stepping

Let $\Delta t > 0$ be the time step size. Let initial conditions $u_{\partial P}^0 \in \mathcal{H}_{\partial P}^{divM}$ be given. Then there exists a unique solution $(u_{\partial P}^n, p^n)$ of Euler-Boussinesq with the following properties:

- ① $(u_{\partial P}^n, p^n)$ conserves global kinetic energy, $E^{kin}(u_{\partial P}^n) = E^{kin}(u_{\partial P}^0)$
- ② $(u_{\partial P}^n, p^n)$ conserves linear momentum $\mathcal{I}(u_{\partial P}^n) = \mathcal{I}(u_{\partial P}^0)$
- ③ $(u_{\partial P}^n, p^n)$ conserves helicity $\mathbf{H}(u_{\partial P}^n) = \mathbf{H}(u_{\partial P}^0)$
- ④ $(u_{\partial P}^n, p^n)$ conserves the potential vorticity $\mathbf{PV}(u_{\partial P}^n) = \mathbf{PV}(u_{\partial P}^0)$

Small Aspect Ratio ϵ

- Thin domain: $\Omega_\epsilon = [-1, 1]^2 \times [-\epsilon, \epsilon]$ transform into $\Omega := [-1, 1]^3$
- Transformation:

$$\mathbf{v}_\epsilon(x, y, z, t) := \mathbf{v}(x, y, \epsilon z, t), \quad p_\epsilon(x, y, z, t) := p(x, y, \epsilon z, t)$$

$$w_\epsilon(x, y, z, t) := \frac{1}{\epsilon} w(x, y, \epsilon z, t).$$

Scaled Euler Equations

$$\partial_t \mathbf{v}_\epsilon + (\mathit{curl} \mathbf{v}_\epsilon \times \mathbf{v}_\epsilon)|_h + \nabla_h \frac{|\mathbf{v}_\epsilon|^2}{2} + \nabla_h p_\epsilon = 0,$$

$$\epsilon^2 \left\{ \partial_t w_\epsilon + (\mathit{curl} \mathbf{v}_\epsilon \times \mathbf{v}_\epsilon)|_v + \partial_z \frac{|\mathbf{v}_\epsilon|^2}{2} \right\} + \partial_z p_\epsilon = 0,$$

$$\mathit{div} \mathbf{v}_\epsilon + \partial_z w_\epsilon = 0.$$

Scaled Euler Equations $\mathbf{v}_\epsilon = (v_1, v_2, v_3)$, $v_i = v_i(x, y, z, t)$

$$\partial_t \mathbf{v}_\epsilon + (\mathit{curl} \mathbf{v}_\epsilon \times \mathbf{v}_\epsilon)|_h + \nabla_h \frac{|\mathbf{v}_\epsilon|^2}{2} + \nabla_h p_\epsilon = 0,$$

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$$\mathit{div} \mathbf{v}_\epsilon + \partial_z w_\epsilon = 0.$$

Hydrostatic Euler Equations $\mathbf{v} = (v_1, v_2)$, $v_i = v_i(x, y, z, t)$

$$\partial_t \mathbf{v} + (\mathit{curl} \mathbf{v} \times \mathbf{v})|_h + \nabla_h \frac{|\mathbf{v}|^2}{2} + \nabla_h p = 0,$$

$$\partial_z p = 0,$$

$$\mathit{div} \mathbf{v}_h + \partial_z w = 0.$$

What happens for $\epsilon \rightarrow 0$?

Discrete Scaled Euler Equations

- $\left\langle \frac{d}{dt} \mathcal{M}_h v^\epsilon + \omega_{\partial \hat{P}} \star U_{\partial P}|_h^{nh} + \mathcal{M}_h \text{grad}_n \left(\frac{E_{kin}^{nh}}{2} + p^\epsilon \right), \phi_h \right\rangle_{\mathcal{H}_F} = 0,$
- $\left\langle \epsilon^2 \left\{ \frac{d}{dt} w^\epsilon + \omega_{\partial \hat{P}} \star U_{\partial P}|_z^{nh} + \mathbf{D}_z \left(\frac{|\mathcal{P} U_{\partial P}^\epsilon|^2}{2} \right) \right\} + \mathbf{D}_z p^\epsilon, \phi_v \right\rangle_{\mathcal{H}_P} = 0,$
- $\mathbf{div}_h \mathcal{M}_h v^\epsilon + \mathbf{div}_v w^\epsilon = 0,$

Discrete Hydrostatic Euler Equations

- $\left\langle \frac{d}{dt} \mathcal{M}_h v + \omega_{\partial \hat{P}} \star U_{\partial P}|_h^{hyd} + \mathcal{M}_h \text{grad}_n \left(p + \frac{E_{kin}^{hyd}}{2} \right), \phi_h \right\rangle_{\mathcal{H}_F} = 0,$
- $\langle \mathcal{P}_z \mathbf{D}_z p, \phi_v \rangle_{\mathcal{H}_P} = 0,$
- $\mathbf{div}_h \mathcal{M}_h v + \mathbf{div}_v w = 0,$

What happens for $\epsilon \rightarrow 0$?

Theorem

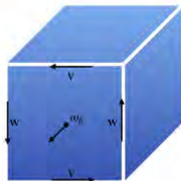
*In the aspect ratio limit, $\epsilon \rightarrow 0$,
the solution $(u_{\partial P}^{nh}, p^{nh})$ of the (nonhydrostatic) Euler equations
converges
to the solution $(u_{\partial P}^{hyd}, p^{hdy})$ of the hydrostatic Euler equations.*

Theorem

In the aspect ratio limit, $\epsilon \rightarrow 0$,
 the solution $(u_{\partial P}^{nh}, p^{nh})$ of the (nonhydrostatic) Euler equations
 converges
 to the solution $(u_{\partial P}^{hyd}, p^{hdy})$ of the hydrostatic Euler equations.

Proof

- Consider equation for the difference $\delta u := u_{\partial P}^{nh} - u_{\partial P}^{hyd}$.
- Analyze difference of nonlinear terms
 - $\omega_{\partial \hat{p}} \star U_{\partial P}|_h^{nh} - \omega_{\partial \hat{p}} \star U_{\partial P}|_h^{hyd}$
 - $w^{nh} \mathcal{P}_h \mathbf{curl}_h U^{nh} - w^{hyd} \mathcal{P}_h \mathbf{D}_z U^{nh}$
 $\sim \mathbf{curl}_h U^{nh} - \mathbf{D}_z U^{nh}$
- Scalar product of difference equation with δu and energy estimate
- Horizontal \mathbf{curl}_h is crucial for estimate



Theorem is discrete version of PDE result by J. Li and E.S. Titi (2019)

Let $\mathcal{G} = \Delta$ be a triangular grid.

$$\begin{aligned}
 \text{Velocity} : & \left\langle \frac{d}{dt} M_h \mathbf{v}, \phi \right\rangle + \left\langle \hat{\mathcal{P}}^T [(f + \omega) \hat{\mathcal{P}} \mathbf{v}], \phi \right\rangle \\
 & + \left\langle \mathcal{P}^T \mathcal{Q} (w \mathbf{D}_z \mathcal{P} \mathbf{v}), \phi \right\rangle + \left\langle \mathcal{M} \mathbf{grad} \left[\frac{|\mathcal{P} \mathbf{v}|_{\mathbb{R}^3}^2}{2} \right], \phi \right\rangle \\
 & + \left\langle \mathcal{P}^T \mathcal{P} \mathbf{grad} (g\eta + p_{hyd}), \phi \right\rangle - \left\langle L \mathbf{v}, \phi \right\rangle = \left\langle \mathcal{F}_v, \phi \right\rangle
 \end{aligned}$$

$$\text{Incompress.} : \mathbf{div}_h \mathcal{M}_h \mathbf{v} + \mathbf{D}_z w = 0$$

$$\text{Free Surface} : \left\langle \frac{\partial \eta}{\partial t}, \psi \right\rangle + \left\langle \mathbf{div} \left[\sum_{k=0}^{k=N_{top}} \mathcal{P}^T (\Delta z_k \mathcal{P} \mathbf{v}_k) \right], \psi \right\rangle = 0$$

$$\begin{aligned}
 \text{Tracer} : & \left\langle \frac{\partial C}{\partial t}, \psi \right\rangle - \left\langle \mathbf{div}^{up} \mathcal{P}^T (C \mathcal{P} \mathbf{v}), \psi \right\rangle + \left\langle LC, \psi \right\rangle \\
 & = \left\langle \mathcal{F}_C, \psi \right\rangle
 \end{aligned}$$

P. K. Formulation of an Unstructured Grid Model for Global Ocean Dynamics (J. Comp. Phys. 339 (2017))

Theorem

- *Let a vertical mixing scheme of PP-type be active.*

Then the semi-discrete hydrostatic Boussinesq ICON-O equations with a free surface have a unique solution, provided the forcing is sufficiently “nice”.

Corollary

The same statement applies if the mesoscale eddy parametrization of Gent-McWilliams-Redi is included and discretized by structure-preserving numerics ^a.

^aP. K. A structure-preserving discretization of ocean parametrizations on unstructured grids (Ocean Modell. (2018))

Now allow the density to vary and to compress

Compressible Euler: Momentum vs Velocity

Momentum: $\partial_t \rho + \operatorname{div}(\rho \mathbf{v}) = 0,$

$$\partial_t(\rho \mathbf{v}) + \rho \operatorname{curl} \mathbf{v} \times \mathbf{v} + \rho \nabla \left(\frac{|\mathbf{v}|^2}{2} + \Phi \right) + \mathbf{v} \operatorname{div}(\rho \mathbf{v}) + \nabla p = 0,$$

$$\partial(\rho e) + \operatorname{div}(\mathbf{v}(\rho e + p)) = 0,$$

Energy: $\rho e := \frac{|\mathbf{v}|^2}{2} + c_V T + \rho \Phi, \quad \text{EOS: } p = \rho RT.$

Velocity: $\partial_t \rho + \operatorname{div}(\rho \mathbf{v}) = 0,$

$$\partial_t \mathbf{v} + \operatorname{curl} \mathbf{v} \times \mathbf{v} + \nabla \left(\frac{|\mathbf{v}|^2}{2} + \Phi \right) + \frac{\nabla p}{\rho} = 0,$$

$$\partial(\rho e) + \operatorname{div}(\mathbf{v}(\rho e + p)) = 0,$$

We use velocity form in analogy with ICON-A.

Similar results for momentum form.

Discrete Compressible Euler

- $$\left\langle \frac{d}{dt} \mathcal{M} u_{\partial P}, \phi \right\rangle_{\mathcal{H}_{\partial P}} + \left\langle \omega_{\partial \hat{P}} \star u_{\partial P} + \mathcal{M} \mathbf{grad} \left(\frac{|\mathcal{P} u_{\partial P}|_{\mathbb{R}^3}^2}{2} \right), \phi \right\rangle_{\mathcal{H}_{\partial P}} + \left\langle \mathcal{P}^T \left(\frac{1}{\rho} \mathcal{P} \mathbf{grad} p \right), \phi \right\rangle_{\mathcal{H}_{\partial P}} = \left\langle \mathcal{M} \mathbf{grad} \Phi, \phi \right\rangle_{\mathcal{H}_{\partial P}},$$
- $$\left\langle \partial_t \rho + \mathbf{div}^{up}(\mathcal{P}^T(\rho \mathcal{P} u_{\partial P})), \psi \right\rangle_{\mathcal{H}_P} = 0,$$
- $$\left\langle \partial_t(\rho \theta) + \mathbf{div}(\mathcal{P}^T(\rho \theta \mathcal{P} u_{\partial P})), \psi \right\rangle_{\mathcal{H}_P} = 0,$$

Theorem (Well-Posedness of Compressible Euler Equations)

Let a time interval $[0, T]$ and initial conditions

- $u_{\partial P}(t=0) = u_0$, and $\theta(t=0) = \theta_0$
- $\rho(t=0) = \rho_0$ with $\rho_0 \geq c > 0$ be given.

Then there exist for $t \in [0, T]$ a unique solution $u_{\partial P}(t)$ of the discrete compressible Euler equations.

We need to assume upwind advection for ρ to avoid vacuum.

Theorem

Solution $u_{\partial P}(t)$ of discrete compressible Euler equations satisfies

- **Energy Conservation:** *The sum of kinetic, potential and internal energy is conserved*

$$\frac{d}{dt}(E^{kin} + E^{pot} + E^{int})(t) = 0, \quad (E^{int} := c_V \rho \theta)$$

- **Helicity Conservation:** *The helicity is conserved*

$$d_t \mathbf{H} = 0.$$

Isentropic Euler Equations with Pressure Equation

$$\partial_t \rho + \mathbf{v} \cdot \nabla \rho + \gamma \rho \operatorname{div}(\mathbf{v}) = 0, \quad (\gamma := \frac{c_v}{c_p})$$

$$\partial_t \mathbf{v} + \operatorname{curl} \mathbf{v} \times \mathbf{v} + \nabla \left(\frac{|\mathbf{v}|^2}{2} + \Phi \right) + \frac{\nabla p}{\rho} = 0.$$

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Discrete Isentropic Euler Equations

- $\langle \partial_t \rho + \operatorname{div}[\mathcal{P}^T(\gamma \rho) \mathcal{P} u_{\partial P}], 1 \rangle_{\mathcal{H}_P} + \langle \gamma' \rho, \operatorname{div} \mathcal{M} u_{\partial P} \rangle_{\mathcal{H}_P} = 0$
- $\langle \frac{d}{dt} \mathcal{M} u_{\partial P}, \phi \rangle_{\mathcal{H}_{\partial P}} + \langle \omega_{\partial \hat{P}} \star u_{\partial P} + \mathcal{M} \operatorname{grad} \left(\frac{|\mathcal{P} u_{\partial P}|_{\mathbb{R}^3}^2}{2} \right), \phi \rangle_{\mathcal{H}_{\partial P}}$
 $+ \langle \mathcal{P}^T \left(\frac{1}{\rho} \mathcal{P} \operatorname{grad} p \right), \phi \rangle_{\mathcal{H}_{\partial P}} = \langle \mathcal{M} \operatorname{grad} \Phi, \phi \rangle_{\mathcal{H}_{\partial P}},$

Theorem

Compressible-Incompressible

- $(u_{\partial P}^\epsilon, p^\epsilon)$ solution of compressible Euler eq.
- $(u_{\partial P}, p)$ solution of incompressible Euler eq.
- well-prepared initial conditions:

$$\mathbf{div} u_{\partial P}^\epsilon(t=0) = \mathcal{O}(\epsilon), \quad p^\epsilon(t=0) = p(t=0) + \mathcal{O}(1)$$

Then solution of compressible equations $(u_{\partial P}^\epsilon, p^\epsilon)$ can be written as

$$u_{\partial P}^\epsilon = \underbrace{u + U}_{\text{slow part}} + \tilde{U} + \mathcal{O}(\epsilon), \quad p^\epsilon = \underbrace{p + P}_{\text{slow part}} + \tilde{P} + \mathcal{O}(\epsilon),$$

where

- (U, P) solution to linearized incompressible Euler
- (\tilde{U}, \tilde{P}) solution to equations of linear acoustics
 $\partial_{tt} P' = \Delta_{\mathcal{M}} P', \quad \mathbf{curl} U' = 0 \quad (\Delta_{\mathcal{M}} u := \mathbf{div} \mathcal{M} \mathbf{grad} u).$

(Proof by analysis of multiscale expansion w.r.t. ϵ)

Discrete version of classical PDE-results from Klainermann-Majda, Kreiss, Schochet. . .

I will discuss the following topics

- i Incompressible Dynamics (*~ ocean*)
- ii Compressible Dynamics (*~ atmosphere*)
- iii Singular Limits (*relation between different equations*)
- iv **Lessons learned**

Observation

- Discrete differential operators & reconstructions mesh-unaware

Consequence: Mesh-Independence

- Results valid for:
triangular \triangle , hexagonal \hexagon and rectangular \square cells
- Results valid for mixed grids $\square\triangle\hexagon\triangle\triangle\square\square$ or Delauny-Voronoi polygons.

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Case of rectangular grids \square

- Discrete differential operators become classical finite differences
- Reconstructions become familiar averages
- **Nonhydrostatic:** MAC method for Navier-Stokes (Harlow-Welch, 1965)
- **Hydrostatic Boussinesq:** same velocity eq. as *NEMO*
 - Nonlinearity conserves 3D-Energy and in 2D energy & enstrophy
 - This is again also valid for triangular and hexagonal meshes

2D Incompressible Euler: $\partial_t \omega + \mathbf{v} \cdot \nabla \omega = 0$

- $\partial_t \Delta \psi + \mathcal{J}(\psi, \Delta \psi) = 0$
- stream function $\mathbf{v} := \nabla^\perp \psi$, $\omega = \Delta \psi$ Jacobian \mathcal{J}

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Arakawa's Jacobian \mathcal{J} conserves energy & enstrophy on quads

$$\mathcal{J}(\psi, \Delta \psi) = \frac{1}{3} \mathcal{J}_1(\psi, \Delta \psi) + \frac{1}{3} \mathcal{J}_2(\psi, \Delta \psi) + \frac{1}{3} \mathcal{J}_3(\psi, \Delta \psi),$$

$$\mathcal{J}_1(\mathbf{p}, \mathbf{q}) := \delta_{2x} p \delta_{2y} q - \delta_{2x} q \delta_{2y} p,$$

$$\mathcal{J}_2(\mathbf{p}, \mathbf{q}) := \delta_{2x} (p \delta_{2y} q) - \delta_{2y} (p \delta_{2x} q),$$

$$\mathcal{J}_3(\mathbf{p}, \mathbf{q}) := \delta_{2y} (q \delta_{2x} p) - \delta_{2x} (q \delta_{2y} p).$$

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Arakawa's Jacobian and $\hat{\mathcal{P}}^\dagger(\omega \hat{\mathcal{P}} \mathbf{v})$

$$\mathcal{J}(\psi, \Delta \psi) = \mathcal{K}(\psi, \Delta \psi)$$

$$\text{with } \mathcal{K}(\psi, \Delta \psi) := \mathbf{curl} \hat{\mathcal{P}}^\dagger(\Delta \psi \hat{\mathcal{P}} \mathbf{grad}_\tau \psi)$$

→ This suggests \mathcal{K} as a generalization of Arakawa's Jacobian to general grids.

C-staggering

Preference of vector invariant nonlinearity: $\text{curl} v \times v + \frac{|v|^2}{2}$

First Way to Instability: Specification of Kinetic Energy $\frac{|v|^2}{2}$?

- C-grid models struggle^a with kinetic energy formulation $|\vec{v}|^2$
- Orthogonal vs non-orthogonal grids
- **Plancherels theorem:** sum of squared components gives vector length if and only if components are from orthonormal basis.
- **Rectangular=Orthogonal** : sum of squared components $|\vec{v}|^2 \sim \sum_{e \in \partial \square} |v_e|^2$ is justified
- **Unstructured=Non-orthogonal:** need to rely on square of reconstructed vector $|\vec{v}|^2 \sim |\mathcal{P}v|^2$
→ This implies a mass matrix \mathcal{M}
- Using sum of squares on unstructured grids creates energy source/sink

^aICOM-A: Zängl, QJRM, 2017, MPAS-A: Skamarock-Klemp, MWR, 2012

Second Way to Instability: Exterior Product $\omega \times v$

- Mixture of vector-invariant and advective form of nonlinearity (*partly vector invariant, partly advective*)
- Prohibits cancelation of fluxes
- Ambiguous nonlinearity impedes energetic consistency and other conservation properties
- Lack of energetic consistency degrades models stability properties

Time Stepping

Fully discrete conservation laws presented here demand implicit time stepping.

Not Negotiable: Algorithmic Essentials

- Clean kinetic energy definition E^{kin} :
 - **Non-orthogonal grids:** reconstruction-based mandatory
 - **Orthogonal grids:** sum of squares or by reconstruction.
- 3D-vector-invariant form of nonlinearity

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Negotiable: Algorithmic Degrees of Freedom

- Reconstructions: different reconstructions can be used
- Vertical coordinates
(we know how to do this)
- Lumping mass matrix in time derivative: short-cut to inverse \mathcal{M}^{-1}
- Higher-Order, upwind-biased reconstructions
- Flux limiters
- Time stepping: alternative time steppings can be used
(implicit used here for theoretical beauty)

The Discrete Hierarchy of Atmosphere-Ocean Equations

